

Asymptotically safe Quantum Gravity

Nonperturbative renormalizability and fractal space-times

Frank Saueressig

Institute for Theoretical Physics & Spinoza Institute

Utrecht University

Rapporteur talk

Quantum Gravity in the Americas III

24th to 26th August 2006

Towards a quantum theory of gravity

- Classical General Relativity:
 - phenomenologically successful at scales $l \gg l_{\text{Pl}}$
(laboratory, solar system, galaxies, ...)
- Quantized General Relativity:
 - theory is perturbatively non-renormalizable:
need infinite number of counter terms
 \iff theory has no predictive power
 - belief that General Relativity is an effective theory:
not fundamental \iff not valid at arbitrary small distances

Towards a quantum theory of gravity

- Classical General Relativity:
 - phenomenologically successful at scales $l \gg l_{\text{Pl}}$
(laboratory, solar system, galaxies, ...)
- Quantized General Relativity:
 - theory is perturbatively non-renormalizable:
need infinite number of counter terms
 \iff theory has no predictive power
 - belief that General Relativity is an effective theory:
not fundamental \iff not valid at arbitrary small distances
- there are fundamental theories which are not perturbatively renormalizable
 - so-called non-perturbatively renormalizable theories
 - same predictive power as a perturbative renormalizable theory

Generalized principles of renormalization (K. Wilson)

- theory is fundamental = infinite UV cutoff limit exists
- infinite cutoff limit requires fixed point in RG flow
 - Gaussian fixed point (GFP) ($u_* = 0$) (e.g. pert. renormalization)
 - non-Gaussian fixed point ($u_* \neq 0$)
- fundamental theory is defined through fixed point
- predictivity linked to how coupling constants approach the fixed point

Weinbergs “asymptotic safety” conjecture

(Weinberg (1979) [hep-th/9702027])

conjecture: Euclidean quantum gravity is asymptotically safe

- RG flow of gravity has non-trivial UV fixed point defining a fundamental theory
- evidence: non-Gaussian fixed point (NGFP) in $2 + \epsilon$ gravity

Weinbergs “asymptotic safety” conjecture

(Weinberg (1979) [hep-th/9702027])

conjecture: Euclidean quantum gravity is asymptotically safe

- RG flow of gravity has non-trivial UV fixed point defining a fundamental theory
- evidence: non-Gaussian fixed point (NGFP) in $2 + \epsilon$ gravity

if NGFP also exists in $d = 4$, it defines a fundamental theory of gravity:

Quantum Einstein Gravity (QEG)

- not clear that fundamental theory is of Einstein-Hilbert form
 \implies approach differs from quantizing General Relativity

Outline

- introduce tool to study non-perturbative RG flows:
exact RG equation (ERGE) for effective average action Γ_k
- evidence for NGFP in $d = 4$ quantum gravity
- consequences of asymptotic safety:
 - fractal structure of QEG space-times
 - dynamical dimensional reduction to $d = 2$

The effective average action Γ_k

- Γ_k is Wilson-type (coarse grained) effective action, based on path integral
- IR cutoff at momentum k^2 :
 - includes all quantum effects with momenta $p^2 > k^2$
 - quantum fluctuations with $p^2 < k^2$ suppressed by mass² = $\mathcal{R}_k(p^2)$
- limits: $k \rightarrow \infty$ = bare/classical action S
 $k \rightarrow 0$ = ordinary effective action Γ
- Γ_k satisfies exact evolution equation

$$k\partial_k\Gamma_k = \frac{1}{2}\text{Tr} [(\delta^2\Gamma_k + \mathcal{R}_k)^{-1}k\partial_k\mathcal{R}_k]$$

- nonperturbative approximation scheme:
“truncating” the space of action functionals

Constructing Γ_k for gravity

- Starting point: path integral for euclidean metrics, $\int \mathcal{D}\gamma \exp[-S_{\text{grav}}[\gamma_{\mu\nu}]]$
- background gauge fixing:
 - decompose quantum metric: $\gamma_{\mu\nu} = \bar{g}_{\mu\nu} + h_{\mu\nu}$
 - add: background gauge fixing $S_{\text{gf}}[h; \bar{g}] + S_{\text{gh}}[h, C, \bar{C}; \bar{g}]$

Constructing Γ_k for gravity

- Starting point: path integral for euclidean metrics, $\int \mathcal{D}\gamma \exp[-S_{\text{grav}}[\gamma_{\mu\nu}]]$
- background gauge fixing:
 - decompose quantum metric: $\gamma_{\mu\nu} = \bar{g}_{\mu\nu} + h_{\mu\nu}$
 - add: background gauge fixing $S_{\text{gf}}[h; \bar{g}] + S_{\text{gh}}[h, C, \bar{C}; \bar{g}]$
- expand $h_{\mu\nu}$ in \bar{D}^2 -eigenmodes and introduce IR cutoff $\Delta_k S$
 - modes with $-\bar{D}^2$ -eigenvalue $> k^2$ are integrated out
 - modes with $-\bar{D}^2$ -eigenvalue $< k^2$ suppressed by mass-term \mathcal{R}_k

Constructing Γ_k for gravity

- Starting point: path integral for euclidean metrics, $\int \mathcal{D}\gamma \exp[-S_{\text{grav}}[\gamma_{\mu\nu}]]$
- background gauge fixing:
 - decompose quantum metric: $\gamma_{\mu\nu} = \bar{g}_{\mu\nu} + h_{\mu\nu}$
 - add: background gauge fixing $S_{\text{gf}}[h; \bar{g}] + S_{\text{gh}}[h, C, \bar{C}; \bar{g}]$
- expand $h_{\mu\nu}$ in \bar{D}^2 -eigenmodes and introduce IR cutoff $\Delta_k S$
 - modes with $-\bar{D}^2$ -eigenvalue $> k^2$ are integrated out
 - modes with $-\bar{D}^2$ -eigenvalue $< k^2$ suppressed by mass-term \mathcal{R}_k
- adding sources:
 - scale dependent generating function for connected Green's functions

$$W_k[\text{sources}; \bar{g}]$$

Constructing Γ_k for gravity

- adding sources:
 - generating function for connected Green's functions $W_k[\text{sources}; \bar{g}]$
- Obtaining the effective average action Γ_k :
 - introduce classical fields: $g_{\mu\nu} = \langle \gamma_{\mu\nu} \rangle_k, \dots$
 - $\Gamma_k [g_{\mu\nu}, \bar{g}_{\mu\nu}, \text{ghosts}] = (\text{modified})$ Legendre transform of W_k

Constructing Γ_k for gravity

- adding sources:
 - generating function for connected Green's functions $W_k[\text{sources}; \bar{g}]$
- Obtaining the effective average action Γ_k :
 - introduce classical fields: $g_{\mu\nu} = \langle \gamma_{\mu\nu} \rangle_k, \dots$
 - $\Gamma_k [g_{\mu\nu}, \bar{g}_{\mu\nu}, \text{ghosts}] = (\text{modified})$ Legendre transform of W_k
- exact RG equation for Γ_k :

$$k\partial_k\Gamma_k = \frac{1}{2}\text{Tr} \left[(\delta^2\Gamma_k + \mathcal{R}_k(-\bar{D}^2))^{-1} k\partial_k\mathcal{R}_k(-\bar{D}^2) \right] + \text{ghost contribution}$$

Constructing Γ_k for gravity

- adding sources:
 - generating function for connected Green's functions $W_k[\text{sources}; \bar{g}]$
- Obtaining the effective average action Γ_k :
 - introduce classical fields: $g_{\mu\nu} = \langle \gamma_{\mu\nu} \rangle_k, \dots$
 - $\Gamma_k [g_{\mu\nu}, \bar{g}_{\mu\nu}, \text{ghosts}] = (\text{modified})$ Legendre transform of W_k

- exact RG equation for Γ_k :

$$k\partial_k\Gamma_k = \frac{1}{2}\text{Tr} \left[(\delta^2\Gamma_k + \mathcal{R}_k(-\bar{D}^2))^{-1} k\partial_k\mathcal{R}_k(-\bar{D}^2) \right] + \text{ghost contribution}$$

- diffeomorphism invariant effective action:

$$\Gamma_k[g] = \Gamma_k[g, \bar{g} = g, \text{ghost} = 0]$$

- theory: specified by RG trajectory:

$$k \mapsto \Gamma_k[g]$$

The Einstein-Hilbert truncation

- Ansatz for Γ_k

$$\Gamma_k = \frac{1}{16\pi G(k)} \int d^d x \sqrt{g} \{-R + 2\Lambda(k)\} + \text{classical gauge fixing \& ghost terms}$$

- running couplings

Newton constant $G(k)$, dimensionless $g(k) = k^{d-2}G(k)$

cosmological constant $\Lambda(k)$, dimensionless $\lambda(k) = \Lambda(k)/k^2$

The Einstein-Hilbert truncation

- Ansatz for Γ_k

$$\Gamma_k = \frac{1}{16\pi G(k)} \int d^d x \sqrt{g} \{-R + 2\Lambda(k)\} + \text{classical gauge fixing \& ghost terms}$$

- running couplings

Newton constant $G(k)$, dimensionless $g(k) = k^{d-2}G(k)$

cosmological constant $\Lambda(k)$, dimensionless $\lambda(k) = \Lambda(k)/k^2$

- substitute ansatz into ERGE and project on subspace:

\implies autonomous β -functions for g, λ

$$k\partial_k g = \beta_g(g, \lambda) , \quad k\partial_k \lambda = \beta_\lambda(g, \lambda).$$

The Einstein-Hilbert truncation

- Ansatz for Γ_k

$$\Gamma_k = \frac{1}{16\pi G(k)} \int d^d x \sqrt{g} \{-R + 2\Lambda(k)\} + \text{classical gauge fixing \& ghost terms}$$

- running couplings

Newton constant $G(k)$, dimensionless $g(k) = k^{d-2}G(k)$

cosmological constant $\Lambda(k)$, dimensionless $\lambda(k) = \Lambda(k)/k^2$

- substitute ansatz into ERGE and project on subspace:

\implies autonomous β -functions for g, λ

$$k\partial_k g = \beta_g(g, \lambda), \quad k\partial_k \lambda = \beta_\lambda(g, \lambda).$$

- β -functions have a non-Gaussian fixed point:

- exists $g^* > 0, \lambda^* > 0$ where $\beta_g(g, \lambda) = \beta_\lambda(g, \lambda) = 0$
- IR repulsive in both g, λ

Evidence for asymptotic safety

Evidence within the exact RG approach:

- within the Einstein-Hilbert approximation:
 - test \mathcal{R}_k -dependence of physical quantities (critical exponents, $g_* \lambda_*, \dots$)
 \implies k-independent in exact theory
- generalized truncation including R^2 -term:
 - NGFP stable under extending truncation space
(O. Lauscher, M. Reuter, hep-th/0110021, hep-th/0206145)

Evidence for asymptotic safety

Evidence within the exact RG approach:

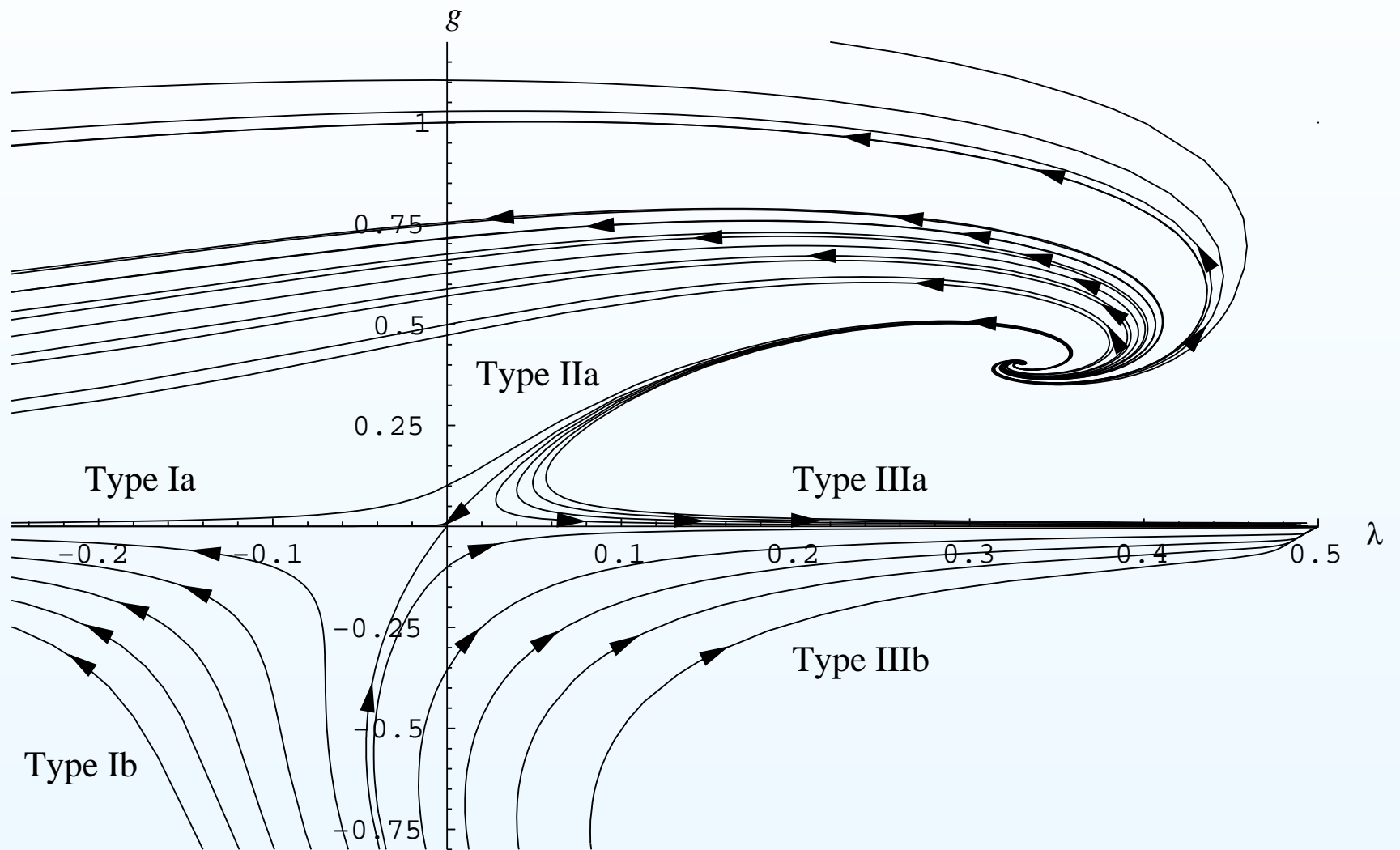
- within the Einstein-Hilbert approximation:
 - test \mathcal{R}_k -dependence of physical quantities (critical exponents, $g_* \lambda_*, \dots$)
 \implies k-independent in exact theory
- generalized truncation including R^2 -term:
 - NGFP stable under extending truncation space
(O. Lauscher, M. Reuter, hep-th/0110021, hep-th/0206145)

NGFP also found via:

- exact path-integral calculation for metrics with 2 Killing vectors
(P. Forgács, M. Niedermaier, hep-th/0207028)
- proper-time RG equation
(A. Bonanno, M. Reuter, hep-th/0410191)

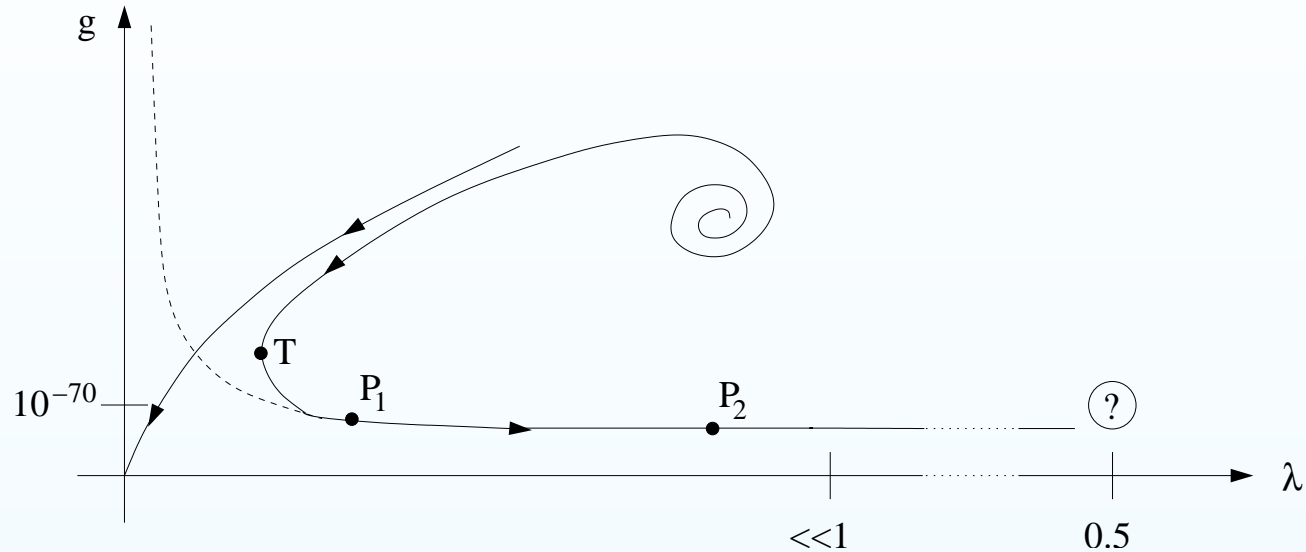
The RG flow of QEG in the Einstein-Hilbert-truncation

(M. Reuter, F. Saueressig, Phys. Rev. D 65 (2002) 065016 [hep-th/0110054])



The RG trajectory realized in Nature

(M. Reuter, H. Weyer, JCAP 0412 (2004) 001 [hep-th/0410119])



- originates at NGFP (quantum regime: $G(k) = k^{-2}g_*$, $\Lambda(k) = k^2\lambda_*$)
- linear regime: oscillations around NGFP
- passing *extremely* close to the GFP
- long classical GR regime (classical regime: $G(k) = \text{const}$, $\Lambda(k) = \text{const}$)
- $\lambda \lesssim 1/2$: strong IR renormalization effects?

Classical versus Quantum space-time

(O. Lauscher and M. Reuter, hep-th/0508202, hep-th/0511260)

RG running of $\Gamma_k[g] \implies$ profound consequences for space-time structure

- Classical General Relativity, S_{EH} :
 - vacuum Einstein equations:

$$R_{\mu\nu} = \Lambda g_{\mu\nu}$$

- solution $g_{\mu\nu}$ = single metric valid at all length scales

Classical versus Quantum space-time

(O. Lauscher and M. Reuter, hep-th/0508202, hep-th/0511260)

RG running of $\Gamma_k[g] \implies$ profound consequences for space-time structure

- Classical General Relativity, S_{EH} :
 - vacuum Einstein equations:

$$R_{\mu\nu} = \Lambda g_{\mu\nu}$$

- solution $g_{\mu\nu}$ = single metric valid at all length scales
- effective action $\Gamma_k[g]$:
 - 1-parameter family of equations of motion:

$$\frac{\delta\Gamma_k[\langle g_{\mu\nu} \rangle_k]}{\delta g_{\mu\nu}} = 0$$

- solution $\langle g_{\mu\nu} \rangle_k$ = metric seen by physical process with momentum k^2
 - proper distances calculated from $\langle g_{\mu\nu} \rangle_k$ depend on k^2
 \implies typical fractal behavior (Coastline of England)

Classical versus Quantum space-time

- effective action $\Gamma_k[g]$:
 - 1-parameter family of equations of motion:

$$\frac{\delta\Gamma_k[\langle g_{\mu\nu}\rangle_k]}{\delta g_{\mu\nu}} = 0$$

- solution $\langle g_{\mu\nu}\rangle_k =$ metric seen by physical process with momentum k^2
 - proper distances calculated from $\langle g_{\mu\nu}\rangle_k$ depend on k^2
 \implies typical fractal behavior (Coastline of England)
- Classical limit:
 - fractal behavior = linked to k -dependence of $\Gamma_k[g_{\mu\nu}]$
 - trajectory realized in nature \iff long classical regime :
 $\Gamma_k[g_{\mu\nu}]$ is k -independent \iff recover classical space-time picture

Relating k to the coarse graining scale $\ell(k)$

- Interpretation: $\langle g_{\mu\nu} \rangle_k$ averaged over proper distances $\ell(k)$

Algorithm to determining $\ell(k)$:

1. Recall: $\Gamma_k[g] = \Gamma_k[g, \bar{g} = g, \text{ghost} = 0]$
 $\implies k^2 = -\bar{D}^2 = -D^2$ is last eigenmode integrated out
2. construct “on-shell”-Laplacians $\Delta(k) = D^2(\langle g_{\mu\nu} \rangle_k)$
3. find eigenfunction $\psi(x)$ with eigenvalue k^2
4. determine typical coordinate distance Δx^μ on which $\psi(x)$ varies
5. calculate proper length $\ell(k)$

$$\ell(k) = \sqrt{\langle g_{\mu\nu} \rangle_k \Delta x^\mu \Delta x^\nu}$$

scaling argument in the quantum regime:

$$\ell \propto k^{-1}$$

Quantum regime: QEG space-times are self-similar fractals

- in the quantum regime ($k^2 \gg m_{\text{Pl}}^2$)

$$\Lambda(k) = \lambda_* k^2, \quad G(k) = g_* k^{2-d}, \quad k \propto 1/\ell$$

- effective field equations in the Einstein-Hilbert approximation:

$$R_{\mu\nu}(\langle g_{\mu\nu} \rangle_k) = \Lambda(k) \langle g_{\mu\nu} \rangle_k$$

- radius of curvature r_c for a typical solution:

$$r_c(k) \propto \Lambda(k)^{-1/2} \propto k^{-1}$$

- Radius of curvature at resolution ℓ :

$$r_c(\ell) \propto \ell$$

\implies zooming into the space-time structure does not change the image

Dynamical dimensional reduction in QEG space-times

Investigate effective graviton propagator

$$\Gamma_k = - \frac{1}{16\pi G(k)} \int d^d x \sqrt{g} R + \dots$$

expanding around flat space

$$\tilde{\mathcal{G}}_k(p) \propto \frac{G(k)}{p^2}$$

classical regime: $G(k) = \text{const}$

$$\tilde{\mathcal{G}} \propto \frac{1}{p^2} \implies \mathcal{G}(x, y) \propto \frac{1}{|x - y|^{d-2}}$$

Dynamical dimensional reduction in QEG space-times

Investigate effective graviton propagator

$$\Gamma_k = - \frac{1}{16\pi G(k)} \int d^d x \sqrt{g} R + \dots$$

expanding around flat space

$$\tilde{\mathcal{G}}_k(p) \propto \frac{G(k)}{p^2}$$

classical regime: $G(k) = \text{const}$

$$\tilde{\mathcal{G}} \propto \frac{1}{p^2} \implies \mathcal{G}(x, y) \propto \frac{1}{|x - y|^{d-2}}$$

fixed point regime: cutoff is $p^2 \implies G(k^2 = p^2) \propto p^{2-d}$

$$\tilde{\mathcal{G}}_k(p) \propto p^{-d} \implies \mathcal{G}(x, y) \propto \ln(\mu|x - y|)$$

Dynamical dimensional reduction in QEG space-times

Investigate effective graviton propagator

$$\Gamma_k = - \frac{1}{16\pi G(k)} \int d^d x \sqrt{g} R + \dots$$

expanding around flat space

$$\tilde{\mathcal{G}}_k(p) \propto \frac{G(k)}{p^2}$$

classical regime: $G(k) = \text{const}$

$$\tilde{\mathcal{G}} \propto \frac{1}{p^2} \implies \mathcal{G}(x, y) \propto \frac{1}{|x - y|^{d-2}}$$

fixed point regime: cutoff is $p^2 \implies G(k^2 = p^2) \propto p^{2-d}$

$$\tilde{\mathcal{G}}_k(p) \propto p^{-d} \implies \mathcal{G}(x, y) \propto \ln(\mu|x - y|)$$

effective graviton propagator reduces dynamically:

macroscopically: $d = 4 \iff$ microscopically: $d = 2$

Spectral dimension for classical manifolds

- diffusion of scalar test particle on Riemannian manifold with metric g

$$\partial_T K_g(x, x'; T) = \Delta_g K_g(x, x'; T)$$

$$\Delta_g \phi \equiv g^{-1/2} \partial_\mu (g^{1/2} g^{\mu\nu} \partial_\nu \phi)$$

- define average return probability

$$P_g(T) \equiv \frac{1}{V} \int d^d x \sqrt{g(x)} K_g(x, x; T)$$

$$= \frac{1}{V} \text{Tr} [\exp(T \Delta_g)]$$

$$= \left(\frac{1}{4\pi T} \right)^{d/2} \sum_{n=0}^{\infty} A_n T^n$$

- asymptotic expansion contains information about space-time dimension

$$d = -2 \frac{d \ln P_g(T)}{d \ln T}$$

Spectral dimension of QEG space-times

- in QEG: metric of manifold is k -dependent
 - \implies diffusion process “with momentum k ” sees metric $\langle g_{\mu\nu} \rangle_k$
 - \implies diffusion equation and return probability will become k -dependent
- Computation of the spectral dimension:
 1. determine k -dependence of $\Delta(k)$
 2. solve the k -dependent heat equation
 3. evaluate “quantum return probability” $P(T)$
 4. obtain spectral dimension

$$\mathcal{D}_s = -2 \frac{d \ln P(T)}{d \ln T}$$

Spectral dimension \mathcal{D}_s of QEG space-times

- Quantum return probability:

$$P(T) = \int \frac{d^d p}{(2\pi)^d} \exp[-p^2 F(p^2) T], \quad F(p^2) = \Lambda(p)/\Lambda(k_0)$$

Limits:

- $T \rightarrow \infty$:
long random walks \iff probe space-time at large distance
 \implies classical regime: $F(p^2)$ with $p^2 \rightarrow 0$
- $T \rightarrow 0$:
short random walks \iff probe space-time at small distance
 \implies fixed point regime: $F(p^2)$ with $p^2 \rightarrow 0$

Spectral dimension \mathcal{D}_s of QEG space-times

- Quantum return probability:

$$P(T) = \int \frac{d^d p}{(2\pi)^d} \exp[-p^2 F(p^2) T] , \quad F(p^2) = \Lambda(p)/\Lambda(k_0)$$

- classical regime: no running, $F(p^2) = 1$:

$$P(T) \propto T^{-d/2} \implies \mathcal{D}_s = d$$

- fixed point regime: $\Lambda(p) \propto p^2 \rightarrow F(p^2) \propto p^2$:

$$P(T) \propto T^{-d/4} \implies \mathcal{D}_s = d/2$$

Spectral dimension \mathcal{D}_s of QEG space-times

- Quantum return probability:

$$P(T) = \int \frac{d^d p}{(2\pi)^d} \exp[-p^2 F(p^2) T], \quad F(p^2) = \Lambda(p)/\Lambda(k_0)$$

- classical regime: no running, $F(p^2) = 1$:

$$P(T) \propto T^{-d/2} \implies \mathcal{D}_s = d$$

- fixed point regime: $\Lambda(p) \propto p^2 \rightarrow F(p^2) \propto p^2$:

$$P(T) \propto T^{-d/4} \implies \mathcal{D}_s = d/2$$

$d = 4$: QEG predicts continuous change of fractal dimension

$\mathcal{D}_s = 4$ macroscopically $\implies \mathcal{D}_s = 2$ microscopically

Spectral dimension \mathcal{D}_s from Causal dynamical triangulations

(J. Ambjørn, et. al. Phys. Rev. Lett. 95 (2005) 171301 [hep-th/0505113])

- CDT \iff Monte Carlo evaluation of Lorentzian Path Integral
- measured \mathcal{D}_s for $d = 4$

$$\mathcal{D}_s(T \rightarrow \infty) = 4.02 \pm 0.1$$

$$\mathcal{D}_s(T \rightarrow 0) = 1.80 \pm 0.25$$

- supports analytic results from exact renormalization group

Summary

- mounting evidence that Euclidean Quantum Gravity is asymptotically safe
- exact renormalization group:
space-time carries a one-parameter family of metrics $\langle g_{\mu\nu} \rangle_k$
describe metric structure at different length-scales
- regimes with non-trivial RG running:
space-time obtains fractal properties
- fractal dimension of space-time continuously changes:
 $\mathcal{D}_s = 4$ at macroscopic distances
 $\mathcal{D}_s = 2$ at microscopic scales